PH4





| Question |  | Marking details | Marks <br> Available |
| :---: | :---: | :---: | :---: |
| 6 | (a) | $\begin{aligned} r & =1.0 \times 10^{8} \mathrm{~m} \text { [unit conversion] (1) } \\ g & =\frac{G M_{\mathrm{E}}}{r^{2}}=\frac{6.67 \times 10^{-11} \times 6.0 \times 10^{24}}{\left(1.0 \times 10^{8}\right)^{2}}(1) \text { [e.c.f. for this mark only] } \\ & =0.04 \mathrm{~N} \mathrm{~kg}^{-1}, \text { Statement "agreement with graph" or equiv (1) } \end{aligned}$ | 2 |
|  | (b) | Moon has a [much] smaller mass than the Earth. [or converse] | 1 |
|  | (c) | $3.45[ \pm 0.05] \times 10^{5} \mathrm{~km} \text { (from graph) (1) }$ <br> No resultant gravitational field [or fields of Earth and Moon equal and opposite] or fields balance at this point. [or equiv](1) | 2 |
|  | (d) | From M to point of intersection / at start $\mathrm{F}_{\text {moon }}>\mathrm{F}_{\text {earth }}$ (1) <br> At point of intersection: $\mathrm{F}_{\text {moon }}=\mathrm{F}_{\text {earth }}$ (1) <br> From point of intersection to earth / at end $\mathrm{F}_{\text {earth }}>\mathrm{F}_{\text {moon }}$ (1) <br> [ -1 for fields rather than forces; -1 not using resultant at least once] | 3 |
|  | (e) | More (1) because gravitational fields of Earth and Moon reinforce [or equiv] and act towards centre of moon opposite to rocket motion. (1) Or [if considering escape from the $\mathrm{E} / \mathrm{M}$ system] Less because of initial greater PE [less negative] due to Earth's field. | 2 |
|  |  |  | [11] |


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| :---: | :---: | :---: | :---: | :---: |
| 7. | (a) |  | $\begin{aligned} & T=1090 \times 24 \times 60 \times 60\left[=9.42 \times 10^{7} \mathrm{~s}\right] \text { [unit conversion] (1) } \\ & r_{\mathrm{s}}=\frac{T v_{\mathrm{s}}}{2 \pi}(1) \text { or equiv e.g. } v=\frac{d}{t} \text { and } d \pi r=6.82 \times 10^{8} \mathrm{~m}(1) \end{aligned}$ | 3 |
|  | (b) | (i) | $T=2 \pi \sqrt{\frac{d^{3}}{G\left(M_{\mathrm{S}}+M_{\mathrm{P}}\right)}}$ (equation selection) (1) [or by impl] <br> $\left(M_{\mathrm{S}} \gg M_{\mathrm{P}}\right)[$ or by impl $] \rightarrow T=2 \pi \sqrt{\frac{d^{3}}{G M_{\mathrm{S}}}}(1)$ <br> $d=\sqrt[3]{\frac{T^{2} G M_{\mathrm{s}}}{4 \pi^{2}}}$ (rearrangement) (1) [or with numbers] <br> Substitution and convincing calculation(1) [to give $\left.=3.21 \times 10^{11} \mathrm{~m}\right]$ | 4 |
|  |  | (ii) | Use of $M_{\mathrm{P}}=\frac{M_{s} r_{s}}{d}$ [in any orientation] or $m_{1} r_{1}=m_{2} r_{2}(1)$ $=\frac{2.2 \times 10^{30} \times 6.8 \times 10^{8}}{3.2 \times 10^{11}}=4.7 \times 10^{27} \mathrm{~kg}(1)$ | 2 |
|  | (c) |  | Find $\Delta \lambda$ in star's spectral lines arising from motion of star / Doppler shift (1) <br> Find velocity of star using $\frac{\Delta \lambda}{\lambda}=\frac{v}{c}$ | 2 |
|  |  |  |  | [11] |

